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The *r*-Hunter–Saxton equation, smooth and singular solutions and their approximation

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Abstract

In this work we introduce the *r*-Hunter–Saxton equation, a generalisation of the Hunter–Saxton equation arising as extremals of an action principle posed in L^r . We characterise solutions to the Cauchy problem, quantifying the blow-up time and studying various symmetry reductions. We construct piecewise linear functions and show that they are weak solutions to the *r*-Hunter–Saxton equation.

Keywords: nonlinear PDEs, singular solutions, Lie symmetries Mathematics Subject Classification numbers: 37K06, 37K58, 35Q53, 37K05.

(Some figures may appear in colour only in the online journal)

1. Introduction

The Hunter–Saxton (HS) equation is a 1 + 1 dimensional, variational, partial differential equation (PDE), originally introduced as a model for the propagation of waves in the director field of a nematic liquid crystal [HS91, HZ95a, HZ95b]. This problem arises in three different forms, related to each other formally through differentiation,

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$$u_t + \left(\frac{u^2}{2}\right)_x = \frac{1}{4} \left(\int_{-\infty}^x - \int_x^\infty\right) u_x^2 \, \mathrm{d}x,\tag{1}$$

$$(u_t + uu_x)_x = \frac{1}{2}u_x^2,$$
 (2)

$$u_{txx} + 2u_{x}u_{xx} + uu_{xxx} = 0, (3)$$

where the subindices denote partial differentiation with respect to the corresponding independent variable. These equations can be considered over a domain Ω that is either the real line, the periodic interval or the interval [a, b] with zero boundary conditions u(a) = u(b) = 0.

The HS equation also has an important geometric interpretation as it describes the geodesic flow on the diffeomorphism group on the domain Ω , with the right-invariant metric defined through the H^1 -inner product

$$\langle f,g\rangle = \int_{\Omega} f_x g_x \,\mathrm{d}x.$$
 (4)

Consider time-parameterised diffeomorphisms $g \in \text{Diff}(\Omega)$ with

$$g_t(x,t) = u(g(x,t),t), \tag{5}$$

then the HS equations can formally be interpreted as extreme points for the action principle

$$\delta \int_0^T l[u] \mathrm{d}t = 0,\tag{6}$$

with

$$l[u] = \int_{\Omega} \frac{1}{2} u_x^2 \, \mathrm{d}x,\tag{7}$$

and endpoint conditions $\delta u(x, 0) = \delta u(x, T) = 0$. This action seeks a geodesic path of diffeomorphisms g(x, t) with $g(x, 0) = g_1(x)$ and $g(x, T) = g_2(x)$ such that the distance functional l[u] is extremised. This problem is the one-dimensional version of problems that arise in computational anatomy [MTY06]. Then, (3) is obtained as the corresponding Euler–Poincaré equation for this action principle after observing that (5) implies that we must have constrained variations $\delta u = w_t + wu_x - uw_x$ for arbitrary w. The conserved energy then arises from Noether's theorem applied to time translation symmetry; we also obtain a Poisson structure for (3) through this route. The Euler–Poincaré derivation is a formal calculation that was made rigorous in [KM03]. A useful connection between these geodesics on the periodic interval and the L^2 -sphere was made in [Len07, Len08] via an explicit mapping; [BBM14] extended this technique to Hunter–Saxton geodesics on the real line.

The integrability of the HS equation was shown in [HZ94] by relating it the Camassa–Holm equation

$$u_t - u_{xxt} + 3uu_x - 2u_x u_{xx} - uu_{xxx} = 0.$$
 (8)

Indeed, it can be shown that (3) arises from a high frequency limit of (8) and, thus, has a bi-Hamiltonian structure [AH09]. Further connections can be made through the geometric interpretation as the Camassa–Holm equation describes geodesic flow with respect to the right-invariant metric [Kou99]

$$\langle f,g\rangle = \int_{\Omega} f_x g_x + fg \, \mathrm{d}x.$$
 (9)

The HS equation has a conserved energy,

$$\frac{\mathrm{d}}{\mathrm{d}t} E_1 \coloneqq \frac{\mathrm{d}}{\mathrm{d}t} \int_{\Omega} u_x^2 \,\mathrm{d}x = 0,\tag{10}$$

which represents one of the two Hamiltonians of the problem, the other being

$$\frac{\mathrm{d}}{\mathrm{d}t} E_2 \coloneqq \frac{\mathrm{d}}{\mathrm{d}t} \int_{\Omega} u u_x^2 \, \mathrm{d}x = 0.$$
(11)

The explicit control of these two functionals ensures some regularity of the solution.

Solutions to this equation can even be written down explicitly. The HS equation given in the forms (1) and (2) has piecewise linear weak solutions that conserve this energy [HZ94]; these solutions can be described by a set of moving points plus the value of u at those points with linear interpolation in between. In [BC05] it was shown how to make sense of these piecewise solutions as weak solutions, and that weak solutions of HS are locally Lipschitz with respect to the initial conditions, making use of techniques from optimal transport. The locality is clear since it is possible to find piecewise linear solutions where two points with different u values collide in finite time, leading to a jump in the solution.

In this paper, we investigate the effect of modifying the distance functional l[u] such that

$$l[u] = \int_{\Omega} \frac{1}{r} |u_x|^r \,\mathrm{d}x,\tag{12}$$

corresponding to the $W^{1,r}(\Omega)$ distance, for general *r* rather than for r = 2. We call the resulting equation the *r*-Hunter–Saxton (*r*-HS) equation, which takes the form

$$\left(|u_x|^{r-2}u_x\right)_{xt} + \left(|u_x|^{r-2}u_x\right)_x u_x + \left(\left(|u_x|^{r-2}u_x\right)_x u\right)_x = 0,$$
(13)

To the author's knowledge, previous work on quantifying the nature of solutions to PDEs arising from energy functionals posed on L^r is limited to the elliptic case. Indeed, a sequence of works initiated by [Aro65] focused on the study of the Euler–Lagrange equations of (12) in the multi-dimensional setting as $r \to \infty$, see [Kat15, Pry18] for an accessible overview. In [KP16] the authors obtained results for the vectorial analogy of (12) and in [KM17, KP16, KP18] a minimisation problem involving second derivatives was examined. In this work, for the first time, we examine the effect modifying the distance functional has on an evolution problem. One of the motivating reasons to modify the distance functional is that a certain amount of regularity is required for a solution to exist to (5) [DGM98], specifically u(x, t) is required to be in $L^1([0, T], W^{1,\infty}(\Omega))$. For the HS equations with r = 2, this is not satisfied in general and we see the loss of the diffeomorphism property when the piecewise-linear solutions blow up as above. With that in mind, we are interested in how the solutions to the resulting PDE behave as $r \to \infty$.

To extend the notion of piecewise linear solutions of the HS equation to the *r*-HS equation we use an optimal control formulation that arises when trying to optimise l[u] such that a set of points moving with *u* are transported from one configuration to another. We shall see that the optimal *u* is then piecewise linear. Following [CH09, GBR11], after eliminating *u*, we obtain a Hamiltonian system for the point locations *q* and their conjugate momenta *p*, with the conserved energy being equivalent to l[u]. For *u* in a finite dimensional state space, it can be shown that after eliminating *p* and *q*, *u* satisfies the Euler–Poincaré equation corresponding to *u*. However, in this particular case, the *r*-HS equation does not make sense because piecewise solutions do not have enough regularity. In this paper, we resolve this situation by showing that

the piecewise linear solutions are weak solutions of another equation whose formal derivative gives the *r*-HS equation.

The rest of this paper is organised as follows. In section 2 we derive the *r*-HS equation. In section 3 we study some fundamental properties of the *r*-HS equation, specifically we characterise smooth solutions to the problem, give a blow-up criteria and give examples of special solutions arising from a symmetry reduction technique. In section 4 we introduce piecewise-linear functions as solutions to an optimal control problem for points on an interval. In section 5 we show that these piecewise-linear functions are weak solutions of an integrated form of the *r*-HS equation. In section 6 we calculate some numerical examples, and finally in 7 we provide a summary and outlook.

2. The r-Hunter–Saxton equation

In this section we introduce the r-HS equation, describe some of its properties.

Definition 2.1. (*r*-Hunter–Saxton equation) The *r*-Hunter–Saxton equation on the interval [a, b] in its most general form is given by

$$\left(|u_x|^{r-2}u_x\right)_{xt} + \left(|u_x|^{r-2}u_x\right)_x u_x + \left(\left(|u_x|^{r-2}u_x\right)_x u\right)_x = 0,\tag{14}$$

with boundary conditions u(a) = u(b) = 0.

One should notice that, as with the (2)-HS equation, the general *r*-HS equation has various equivalent forms.

Proposition 2.2. (*Different forms of* (14)) *The following are formally equivalent formulations of* (14):

$$\left(|u_x|^{r-2}u_x\right)_{xt} + \left(|u_x|^{r-2}u_x\right)_x u_x + \left(\left(|u_x|^{r-2}u_x\right)_x u\right)_x = 0,$$
(15)

$$|u_x|^{r-2}u_{xt} + \frac{1}{r}|u_x|^r + |u_x|^{r-2}u_{xx}u = c(t),$$
(16)

and

$$\left(|u_x|^{r-2}u_x\right)_t + \left(|u_x|^{r-2}uu_x\right)_x = \frac{1}{r}|u_x|^r + c(t),\tag{17}$$

for some c(t) that ensures u(b) = 0. Notice that setting r = 2, in (17) yields exactly (2).

Proof. Firstly note that

$$\frac{\partial}{\partial x} \left(\left| u_x \right|^r \right) = r \left| u_x \right|^{r-2} u_x u_{xx},\tag{18}$$

and hence

$$\frac{\partial}{\partial x} \left(|u_x|^{r-2} u_x \right) = (r-2) |u_x|^{r-4} u_x^2 u_{xx} + |u_x|^{r-2} u_{xx},$$

= $(r-1) |u_x|^{r-2} u_{xx}.$ (19)

Similarly,

$$\frac{\partial}{\partial t} \left(\left| u_x \right|^{r-2} u_x \right) = (r-1) \left| u_x \right|^{r-2} u_{xt}.$$
(20)
7019

Making use of these we see that

$$0 = \left(|u_{x}|^{r-2}u_{x}\right)_{tx} + \left(|u_{x}|^{r-2}u_{x}\right)_{x}u_{x} + \left(\left(|u_{x}|^{r-2}u_{x}\right)_{x}u\right)_{x},$$

$$= (r-1)\left(|u_{x}|^{r-2}u_{xt}\right)_{x} + (r-1)|u_{x}|^{r-2}u_{xx}u_{x} + (r-1)\left(|u_{x}|^{r-2}u_{xx}u\right)_{x},$$

$$= (r-1)\left(|u_{x}|^{r-2}u_{xt}\right)_{x} + \frac{(r-1)}{r}\left(|u_{x}|^{r}\right)_{x} + (r-1)\left(|u_{x}|^{r-2}u_{xx}u\right)_{x}.$$
 (21)

Hence (15) is the formal derivative of (16).

To see (17) note that

$$\frac{\partial}{\partial x} \left(|u_x|^{r-2} u u_x \right) = (r-2) |u_x|^{r-4} u_x^2 u_{xx} u + |u_x|^{r-2} u_x^2 + |u_x|^{r-2} u u_{xx},$$
$$= (r-1) |u_x|^{r-2} u u_{xx} + |u_x|^r.$$
(22)

Hence

$$c(t) = \left(|u_x|^{r-2}u_x\right)_t + \frac{(r-1)}{r}|u_x|^r + (r-1)|u_x|^{r-2}u_{xx}u,$$

= $\left(|u_x|^{r-2}u_x\right)_t - \frac{1}{r}|u_x|^r + \left(|u_x|^{r-2}uu_x\right)_x,$ (23)

as required.

The equations can also be defined on the real line, or with periodic boundary conditions, but we concentrate on the boundary value problem in this paper for simplicity.

Next we show that the *r*-Hunter–Saxton equation emerges from Hamilton's principle for a $W^{1,r}$ Lagrangian.

Proposition 2.3 (Euler–Poincaré equation). *The r-Hunter–Saxton* (14) *is the Euler–Poincaré equation for the Lagrangian*

$$l[u] = \int_{\Omega} \frac{1}{r} |u_x|^r \,\mathrm{d}x. \tag{24}$$

Proof. We have

$$\delta S = \delta \int_{a}^{T} \int_{a}^{b} \frac{1}{r} |u_{x}|^{r} \, \mathrm{d}x \, \mathrm{d}t, \tag{25}$$

$$= \int_{0}^{T} \int_{a}^{b} |u_{x}|^{r-2} u_{x} \delta u_{x} \, \mathrm{d}x \, \mathrm{d}t, \tag{26}$$

$$= \int_0^T \int_a^b -\left(|u_x|^{r-2}u_x\right)_x \delta u \, \mathrm{d}x \, \mathrm{d}t, \tag{27}$$

through an integration by parts making use of the boundary conditions u(a, t) = u(b, t) = 0. Following [HMR98], (5) implies that

$$\delta u = w_t - w u_x + u w_x, \tag{28}$$

with w(x, 0) = w(x, T) = 0. Substituting (28) into (25), we obtain

$$0 = \delta S = \int_0^T \int_a^b -\left(|u_x|^{r-2}u_x\right)_x (w_t - wu_x + uw_x) \,\mathrm{d}x \,\,\mathrm{d}t,\tag{29}$$

$$= \int_{0}^{T} \int_{a}^{b} \left(\left(|u_{x}|^{r-2} u_{x} \right)_{xt} + \left(|u_{x}|^{r-2} u_{x} \right)_{x} u_{x} + \left(\left(|u_{x}|^{r-2} u_{x} \right)_{x} u \right)_{x} \right) w \, \mathrm{d}x \, \mathrm{d}t, \quad (30)$$

which is satisfied by solutions of (14) for arbitrary w, as required.

Later we shall make use of the following weak form of (17).

Definition 2.4 (Weak integrated *r*-HS equation). Let $W_0^{1,r}(a, b)$ be the space of functions with

$$\int_{a}^{b} \left| u \right|^{r} + \left| u_{x} \right|^{r} \, \mathrm{d}x < \infty, \tag{31}$$

that satisfy the boundary conditions u(a) = u(b) = 0. Then, $u \in W_0^{1,r}(a, b)$ is a weak solution of (17) if it satisfies

$$\int_{a}^{b} \left(\left(|u_{x}|^{r-2} u_{x} \right)_{t} - \frac{1}{r} |u_{x}|^{r} \right) \phi - u |u_{x}|^{r-2} u_{x} \phi_{x} \, \mathrm{d}x = c(t) \int_{a}^{b} \phi \, \mathrm{d}x, \tag{32}$$

for all test functions $\phi \in W_0^{1,r}(a,b)$.

3. Symmetries and classical solutions

In this section we examine some symmetries and characterise classical solutions and their existence time for the *r*-HS equation. We take inspiration from the arguments in [[HS91], section 3] and show that, remarkably, many of the properties shown for the r = 2 case generalise for arbitrary *r*. We begin by examining a characteristic reduction. To that end, throughout this section we will assume $u \in C^2(\mathbb{R})$ be a classical solution of the Cauchy problem

$$(|u_x|^{r-2}u_x)_t + (|u_x|^{r-2}uu_x)_x = \frac{1}{r}|u_x|^r,$$

$$u(x,0) = u_0(x).$$
(33)

Theorem 3.1 (Smooth solution characterisation). *Every smooth solution of* (33) *can be written implicitly as*

$$u = H'(t) + \sum_{k=1}^{r} kC_k t^{k-1} G_k(\xi),$$

$$x = H(t) + \xi + \sum_{k=1}^{r} C_k t^k G_k(\xi).$$
(34)

where $H \in C^1(\mathbb{R})$ is any function with H(0) = H'(0) = 0 and, for $k \in \mathbb{N}$, $G_k \in C^2(\mathbb{R})$ is defined by requiring

$$G'_k(\xi) = u'_0(\xi)^k.$$
(35)

The constants C_k are defined as

$$C_k \coloneqq \frac{r!}{r^k k! (r-k)!}.$$
(36)

Proof. Let ξ denote a characteristic curve with $U(\xi, t) = u(X(\xi, t), t)$ and X satisfying the initial value problem

$$X_t(\xi, t) = U(\xi, t),$$

 $X(\xi, 0) = \xi.$
(37)

Further, let $V(\xi, t) = X_{\xi}(\xi, t)$. It can then be verified that V satisfies

$$V_t = u_X X_{\xi},$$

$$\left(|V_t|^{r-2} V_t \right)_t = (r-1) \left(|X_{\xi} u_X|^{r-2} X_{\xi} \left(u_{XX} u + u_{Xt} + u_X^2 \right) \right),$$
(38)

and hence solves the second order initial value problem

$$V(|V_t|^{r-2}V_t)_t = \frac{(r-1)^2}{r} |V_t|^r,$$

$$V(\xi, 0) = 1,$$

$$V_t(\xi, 0) = u'_0(\xi).$$

(39)

This can then be solved to show that

$$V = \left(\frac{tu_0'(\xi)}{r} + 1\right)^r.$$
(40)

This should be compared with the r = 2 case found in [[HS91], section 3]. The final result follows from solving the system

$$U(\xi, t) = X_t(\xi, t),$$

$$X_{\xi}(\xi, t) = \left(\frac{tu'_0(\xi)}{r} + 1\right)^r,$$
(41)

for U to show (34).

It is also clear that any functions satisfying (34) are indeed solutions of (33) so long as the relation for the characteristic curve is invertible. In view of this, it is expected that smooth solutions break in finite time.

Corollary 3.2 (Existence and uniqueness). *Suppose the Cauchy problem* (33) *is coupled with an asymptotic boundary condition that*

$$\lim_{x \to \infty} u(x,t) = 0 \tag{42}$$

and let the initial condition $u_0 \in W^{1,r}(\mathbb{R})$ decay asymptotically, that is

$$\lim_{x \to \infty} u_0(x,t) = 0. \tag{43}$$

Then, there is a T > 0 such that for $t \in (0, T)$ there is a unique smooth solution of the Cauchy problem coupled to the asymptotic boundary condition (42) given by (34), $H(t) \equiv 0$ and

$$G_k(\xi) = -\int_{\xi}^{\infty} u'_0(s)^k \mathrm{d}s \quad for \ k \in [1, r].$$
(44)

Theorem 3.3 (Blow up time). Let $T = \frac{r}{\sup_{x \in \Omega} - u'_0(\varepsilon)}$ then, if u_0 is not monotonically increasing, (33) has a smooth solution for all $t \in (0, T)$ and $\sup |u_x| \to \infty$ as $t \to T$.

Proof. The implicit function theorem guarantees that as long as $X_{\xi} \neq 0$ there is a smooth solution of (33). Given

$$X_{\xi} = \left(\frac{tu_0'(\xi)}{r} + 1\right)^r,\tag{45}$$

we see that $X_{\xi} = 0$ if and only if

$$t = \frac{r}{-u_0'(\xi)}.$$
(46)

So $T = \frac{r}{\sup -u'_0(\xi)}$ as required.

Remark 3.4 (Relating to Burgers' equation). At this point we note that the 1-HS equation is formally equivalent to the inviscid Burgers' equation. Indeed, differentiating

$$u_t + uu_x = 0 \tag{47}$$

in space then dividing through by $|u_x|$ yields (16) with r = 1. This means that for $1 \le r \le 2$, the *r*-HS equations can be viewed as an interpolation between Burgers' equation and the classic HS equation.

Notice the maximal smooth solution time of (33) increases linearly as r increases. In fact, it is exactly r-times the shock time for the inviscid Burgers' equation. In particular, as $r \to \infty$, the blow up time $T \to \infty$. We will return to this point later in this work.

3.1. Symmetries

In this section we find all possible exact solutions that result from inspecting the symmetries of the equation. For the rest of this section, for simplicity of computation, we make the assumption that $r \ge 2$ is even. A Lie point symmetry of equation (33) is a flow

$$\left(\widetilde{x},\widetilde{t},\widetilde{u}\right) = \left(e^{\epsilon X}x, e^{\epsilon X}t, e^{\epsilon X}u\right),\tag{48}$$

generated by a vector field

$$X = \xi^{1}(x, t, u)\frac{\partial}{\partial x} + \xi^{2}(x, t, u)\frac{\partial}{\partial t} + \eta(x, t, u)\frac{\partial}{\partial u},$$
(49)

such that $\tilde{u}(\tilde{x}, \tilde{y})$ is a solution of (33) whenever u(x, y) is a solution of (33). As usual, we denote by $e^{\epsilon X}$ the *Lie series* $\sum_{k=0}^{\infty} \frac{\epsilon^k}{k!} X^k$ with $X^k = XX^{k-1}$ and $X^0 = 1$. To find the symmetries of the *r*-HS equation we are required to solve the infinitesimal invari-

To find the symmetries of the *r*-HS equation we are required to solve the infinitesimal invariance condition for the vector field (49). To do this we use the prolongation of X [BA08, Olv93, Ste89]. The infinitesimal symmetry condition decomposes to a large overdetermined system of linear PDEs for ξ^1 , ξ^2 and η known as *determining equations*. The following three propositions give the overdetermined system, the general form of the determining equations and the Lie algebra generators. These results were proven symbolically using the SYM package

[DT04, DT06]. This procedure is described in further detail in the case of the *p*-Laplacian in [PP19].

Proposition 3.5 (Infinitesimal invariance). *The infinitesimal invariance condition is equivalent to the following system of 14 equations:*

$$\xi_{xu}^2 = \xi_x^2 = \xi_u^2 = \xi_{uu}^2 = 0, \tag{50}$$

$$\eta_{xt} + u\eta_{uu} = \eta_{xu} - \xi_{xt}^2 - u\xi_{xx}^2 = \xi_u^1 - u\xi_u^2 = \xi_u^1 - 2u\xi_u^2 = ru\xi_{uu}^1 - \xi_u^1 = 0,$$

$$-\xi_u^1 + u(\xi_u^2 + r(\xi_{uu}^1 + u\xi_{uu}^2)) = \eta - \xi_t^1 + u(\xi_t^2 - \xi_x^1 + 2u\xi_x^2)$$
(51)

$$= r \eta_{uu} - r \xi_{uu}^2 + r (\xi_{uu}^1 + u \xi_{uu}^2)) = \eta - \xi_t^1 + u (\xi_t^2 - \xi_x^1 + 2u \xi_x^2)$$
$$= r \eta_{uu} - r \xi_{uu}^2 - 2\xi_x^2 - r \xi_{uu}^1 - 2r u \xi_{uu}^2 = 0,$$
(52)

$$r\eta_{tu} + 2\eta_x + r\left(2u\eta_{xu} - \xi_{xt}^1 - u\xi_{xx}^1\right)$$

= $-\eta + u\eta_u + ru^2\eta_{uu} + \xi_t^1 - ru\xi_{tu}^1 - 2ru^2\xi_{xu}^1 = 0.$ (53)

Solutions of the overdetermined system of linear PDEs (50)–(53) will yield the algebra of the symmetry generators (49) of the r-HS equation.

Given (50)–(53) form an overdetermined system of linear partial differential equations it is possible that they only admit the trivial solution $\xi^1 = \xi^2 = \eta = 0$. This would imply that the only Lie symmetry of the *r*-HS equation is the identity transformation. In what follows we will see that this is not the case. We are able to obtain the Lie algebra for the symmetry generators the *r*-HS equation and thus, using the Lie series, derive the groups of Lie point symmetries.

Proposition 3.6 (Determining equations). *The general solution of the determining equations* (50)–(53) *is given by*

$$\xi^{1} = c_{1}t + c_{2}xt + c_{3} + c_{4}x, \quad \xi^{2} = \frac{c_{2}t^{2}}{r} + c_{4}t - c_{5}t + c_{6},$$

$$\eta = c_{1} + c_{2}\left(tu - \frac{2tu}{r} + x\right) + c_{5}u,$$
(54)

where c_i , i = 1, ..., 6 are arbitrary real constants.

Proposition 3.7 (Lie algebra generators). It follows that the solution (54) defines a six dimensional Lie algebra of generators where a basis is formed by the following vector fields

$$X_1 = \frac{\partial}{\partial x}, \quad X_2 = \frac{\partial}{\partial t}, \quad X_3 = \frac{\partial}{\partial u} + t\frac{\partial}{\partial x}, \quad X_4 = t\frac{\partial}{\partial t} + x\frac{\partial}{\partial x},$$
 (55)

$$X_5 = u\frac{\partial}{\partial u} - t\frac{\partial}{\partial t}, \quad X_6 = \frac{t^2}{r}\frac{\partial}{\partial t} + \left(tu - \frac{2tu}{r} + x\right)\frac{\partial}{\partial u} + tx\frac{\partial}{\partial x}.$$
 (56)

3.2. Invariant solutions through symmetry reductions

We now state solutions that occur through symmetry reductions of (33) to ODEs by means of the algebra generators given in proposition 3.7. We consider each generator separately and examine some examples of solutions from each.

 X_1 . To begin notice that solutions of (33) that are invariant under the symmetry generated by X_1 are of the form u = f(t), which immediately yields $u \equiv \text{const}$ as a trivial solution prescribed by the initial condition.

 X_2 . Solutions invariant under the symmetry generated by X_2 are of the form u = f(x). The reduced equation is given by the following ODE:

$$(f_x)^r (f_x^2 + rf f_{xx}) = 0. (57)$$

This, in turn shows that either $f \equiv \text{const}$ or $f_x^2 + rf f_{xx} = 0$ and must take the form

$$f(x) = c_2(x + rx - rc_1)^{r/(1+r)},$$
(58)

for arbitrary constants c_1, c_2 .

*X*₃. Solutions invariant under the symmetry generated by *X*₃ are of the form $u = xt^{-1} + f(t)$ with *f* prescribed by the initial condition.

 X_4 . The quantities u and $\xi = tx^{-1}$ are algebraic invariants of the Lie group generated by X_4 . Assume $u = f(\xi)$ for non-constant f, then we obtain the reduced equation as the following ODE:

$$rf'(\xi)(2\xi f(\xi) - 1) + \xi^2 f'(\xi)^2 + r\xi f''(\xi)(\xi f(\xi) - 1) = 0.$$
(59)

The general solution of this ODE is not known.

 X_5 . Solutions invariant under the symmetry generated by X_5 are of the form $u = f(x)t^{-1}$, for non-constant *f*. The reduced ODE is given by:

$$rf'(x) - f'(x)^{2} - rf(x)f''(x) = 0,$$
(60)

whose solution is an inverse hypergeometric function, that is

$$f^{-1}(x) = \frac{{}_{2}F_{1}\left(1, -r, 1-r, \frac{c(x+c_{2})^{-1/r}}{r}\right)(x+c_{2})}{r},$$
(61)

for constants c_1, c_2 .

*X*₆. The quantities *u* and $\xi = tx^{-1/r}$ are algebraic invariants of the Lie group generated by *X*₆. With $u = f(\xi)$ we may derive the following reduced ODE:

$$-\xi f(\xi) \left(\left(-r^2 + 3r + 4 \right) f'(\xi) + \xi r f''(\xi) \right) + \left(r^2 - 4 \right) f(\xi)^2 - \xi^2 f'(\xi)^2 = 0,$$
(62)

which has closed form solution

$$f(\xi) = c_2 \exp\left(\frac{2r \log(c_1 + \xi^r) + (-2r - 4)\log(\xi)}{2(r+1)}\right).$$
(63)

4. Piecewise linear solutions

In this section we consider an optimal control formulation that allows us to quantify piecewise linear solutions to the problem. We show that the computation of these solutions requires the solution of a nonlinear system and make various comparisons to the 2-HS equation. We leave the formal interpretation of these solutions to the next section.

Definition 4.1 (Optimal control problem). Let $Q_1(t), \ldots, Q_N(t)$ represent a moving set of points on the interval [a, b]. The $W^{1,r}$ optimal control problem is to find $Q_1(t), \ldots, Q_N(t)$ and $u \in W^{1,r}(a, b)$ such that

$$\int_0^T \int_a^b \frac{1}{r} |u_x|^r \,\mathrm{d}x \,\mathrm{d}t \tag{64}$$

is minimised, subject to the constraints

$$\dot{Q}_{i}(t) = u(Q_{i}(t), t),$$

$$Q_{i}(0) = Q_{i}^{A},$$

$$Q_{i}(T) = Q_{i}^{B},$$
(65)

for i = 1, ..., N.

This problem has the following variational formulation.

Definition 4.2 (Clebsch variational principle). Let $Q_1(t), \ldots, Q_N(t)$ be a set of points on the interval [a, b] with Lagrange multipliers P_1, \ldots, P_N , and suppose $u \in W^{1,r}(a, b)$. The Clebsch variational principle corresponding to the optimal control problem in definition 4.1 is

$$\delta S[u, P, Q] = 0, \quad S = \int_{t=0}^{T} l[u] + \sum_{i=1}^{n} P_i(\dot{Q}_i - u(Q_i)) dt, Q_i(0) = Q_i^{A},$$

$$Q_i(T) = Q_i^{B}, \quad i = 1, \dots, N.$$
(66)

This type of variational principle takes its name from variational principles of this form that can be used to derive equations of fluid dynamics, where the Lagrange multipliers (Clebsch variables) enforce the dynamics of transported quantities.

Remark 4.3. In [CH09, GBR11] these variational principles were considered in a general form (to which the name 'Clebsch' was extended) where the Lagrange multipliers enforce dynamics

$$\dot{Q} = L_u Q, \tag{67}$$

where Q is a curve on a manifold \mathcal{M} , u is a curve on a Lie algebra \mathcal{X} , and L_u represents a Lie algebra action on \mathcal{M} . In this case, it was shown that P and Q can be eliminated via the closure of the corresponding Lie algebra bracket. Then, u solves the corresponding Euler–Poincaré equation. In our case, we have $(L_uQ)_i = u(Q_i, t)$, which is a Lie algebra action of smooth vector fields on [a, b], which have the Lie bracket $[u, v] = u_x v - uv_x$. However, $W^{1,r}(a, b)$ is not closed under this bracket, and we see that we do not have enough regularity to complete this argument. However, as we shall find, it is still possible to understand the evolution of u as a weak solution of the corresponding Euler–Poincaré equation (the r-Hunter–Saxton equation, in this case).

Lemma 4.4. The optimal u to the variational problem (66) are given by piecewise linear functions, whose jumps occurring at $x = Q_i$, i = 1, ..., N, and jump condition

$$-\left[|u_x|^{r-2}u_x\right]_{Q_i^-}^{Q_i^+} = P_i.$$
(68)

Proof. The Euler–Lagrange equation corresponding to *u* is given by

$$\int_{0}^{1} |u_{x}|^{r-2} u_{x} \delta u_{x} \, \mathrm{d}x = \sum_{i=1}^{n} P_{i} \delta u(Q_{i}), \tag{69}$$

which is, in turn, the weak form of the r-Laplace equation

$$-\left(|u_x|^{r-2}u_x\right)_x = \sum_{i=1}^n P_i \delta(x - Q_i),$$
(70)

where $\delta(x)$ is the Dirac measure. This means that *u* is piecewise linear, with jump conditions given by (68). Notice that this means that the term $(|u_x|^{r-2}u_x)_x u_x$ in (14) does not make sense, since at $x = Q_i$ it is the product of a Dirac measure with a function that has a jump at the same location. Later we shall see that this problem is resolved since *u* solves (32) nevertheless.

Remark 4.5. Equation (70) defines an infinitesimally equivariant momentum map from $(P, Q) \in T^*([a, b]^n)$ to the dual space $\mathcal{X}([a, b])$ containing the momentum $m = -(|u_x|^{r-2}u_x)_x$. The equivariance explains why it is possible to eliminate *P* and *Q* and obtain an equation for *u* alone. This was explored in the context of Peakon solutions of the Camassa–Holm equation in [HM05], where the infinitesimal equivariance is proved.

Lemma 4.6. The piecewise solution can be characterised through the points $\hat{u}_i = u(Q_i)$, i = 1, ..., n with $\hat{u}_0 = 0$, $\hat{u}_{n+1} = 0$. Then, the set of coefficients $\{\hat{u}_i\}_{i=0}^{n+1}$ solves the difference equation

$$P_{i} = -\left|\frac{\hat{u}_{i+1} - \hat{u}_{i}}{Q_{i+1} - Q_{i}}\right|^{r-2} \left(\frac{\hat{u}_{i+1} - \hat{u}_{i}}{Q_{i+1} - Q_{i}}\right) + \left|\frac{\hat{u}_{i} - \hat{u}_{i-1}}{Q_{i} - Q_{i-1}}\right|^{r-2} \times \left(\frac{\hat{u}_{i} - \hat{u}_{i-1}}{Q_{i} - Q_{i-1}}\right) \quad for \ i = 1, \dots, n,$$

$$\hat{u}_{0} = \hat{u}_{n+1} = 0.$$
(71)

Proof. The piecewise linear solution has piecewise constant derivative,

$$u_x = \frac{\hat{u}_{i+1} - \hat{u}_i}{Q_{i+1} - Q_i}, \quad i = 1, \dots, n,$$
(72)

and substitution into (68) gives the result.

Lemma 4.7. Let $Q_{i+1} > Q_i$ for i = 0, ..., n. Then, equation (71) has a unique solution. **Proof.** Let

$$\hat{l}(\hat{u}, Q) = \sum_{i=0}^{n} \frac{|\hat{u}_{i+1} - \hat{u}_i|^r}{r|Q_{i+1} - Q_i|^{r-2} (Q_{i+1} - Q_i)}.$$
(73)

The solution $\{\hat{u}_i\}_{i=0}^{n+1}$ is then the minimiser of

$$\hat{l}(\hat{u}, Q) - \sum_{i=1}^{n} P_i \hat{u}_i.$$
 (74)

Writing $\Delta \hat{u}_i = \hat{u}_{i+1} - \hat{u}_i$, i = 0, ..., n, we see that $\{\Delta \hat{u}_i\}_{i=0}^{n+1}$ minimises the function

$$\sum_{i=0}^{n} \frac{|\Delta \hat{u}_{i}|^{r}}{r|Q_{i+1} - Q_{i}|^{r-2} (Q_{i+1} - Q_{i})} - \sum_{i=1}^{n} P_{i} \sum_{j=0}^{i} (\Delta \hat{u}_{i}),$$
(75)

subject to the constraint $\sum_{i=1}^{n} \Delta \hat{u}_i = 0$. The function is the sum of convex and linear functions, which is then also convex. Since the constraint defines a linear subspace, we are minimising a convex function over a finite dimensional vector space, which has a unique solution.

The solution does not have a closed form analytic expression but can be solved numerically using Newton's method. We denote the solution operator $\hat{u} = U(P, Q)$.

Theorem 4.8. The equations for P and Q satisfy

$$Q_i = \hat{u}_i,\tag{76}$$

$$\dot{P}_{i} = \frac{r-1}{r} \left(\left| \frac{\hat{u}_{i} - \hat{u}_{i-1}}{Q_{i} - Q_{i-1}} \right|^{r} - \left| \frac{\hat{u}_{i+1} - \hat{u}_{i}}{Q_{i+1} - Q_{i}} \right|^{r} \right),$$
(77)

$$\hat{u} = U(P, Q),\tag{78}$$

with the conserved energy

$$E[P,Q] = \hat{l}(U(P,Q),Q).$$
(79)

Proof. We know from [CH09, GBR11] that the equations for P and Q are canonical Hamiltonian with Hamiltonian function given by Legendre transform

$$H(P,Q) = \sum_{j} P_{j} \hat{u}_{j} - \hat{l}(\hat{u},Q), \quad \hat{u} = U(P,Q),$$
(80)

since $\hat{l}(\hat{u}, Q)$ is equal to l(u) when u is the piecewise linear function interpolating \hat{u} . The equation for Q is then

$$\dot{Q}_{i} = \frac{\partial}{\partial P_{i}} H(P,Q) = \hat{u}_{i} + \sum_{j} \underbrace{\left(P_{j} - \frac{\partial}{\partial \hat{u}_{j}} \hat{l}(\hat{u},Q)\right)}_{=0} \frac{\partial}{\partial P_{i}} U_{j}(P,Q) = \hat{u}_{i}, \quad (81)$$

as expected, after noting that the bracket vanishes since \hat{u} satisfies (71). After noting that

$$\frac{\partial}{\partial Q_{i}}\hat{l}(\hat{u},Q) = -\frac{r-1}{r} \left(\left| \frac{\hat{u}_{i} - \hat{u}_{i-1}}{Q_{i} - Q_{i-1}} \right|^{r} - \left| \frac{\hat{u}_{i+1} - \hat{u}_{i}}{Q_{i+1} - Q_{i}} \right|^{r} \right),$$
(82)

we see the equation for P is given by

$$\dot{P}_i = -\frac{\partial H}{\partial Q_i},\tag{83}$$

$$=\sum_{j}\underbrace{\left(P_{j}-\frac{\partial}{\partial\hat{u}_{j}}\hat{l}(\hat{u},Q)\right)}_{=0}\frac{\partial}{\partial Q_{i}}U_{j}(P,Q)-\frac{\partial}{\partial Q_{i}}\hat{l}(\hat{u},Q),\tag{84}$$

$$=\frac{r-1}{r}\left(\left|\frac{\hat{u}_{i}-\hat{u}_{i-1}}{Q_{i}-Q_{i-1}}\right|^{r}-\left|\frac{\hat{u}_{i+1}-\hat{u}_{i}}{Q_{i+1}-Q_{i}}\right|^{r}\right),$$
(85)

as required.

Hence, solving the nonlinear equation (71) for \hat{u} allows us to calculate both \dot{P} and \dot{Q} and numerically integrate the equations.

Note that when r = 2, the P equation specialises to

$$\dot{P}_{i} = \frac{1}{2} \left(\frac{\hat{u}_{i} - \hat{u}_{i-1}}{Q_{i} - Q_{i-1}} \right)^{2} - \frac{1}{2} \left(\frac{\hat{u}_{i+1} - \hat{u}_{i}}{Q_{i+1} - Q_{i}} \right)^{2},\tag{86}$$

$$= \frac{1}{2} \left(-\left(\frac{\hat{u}_{i+1} - \hat{u}_i}{Q_{i+1} - Q_i}\right) + \left(\frac{\hat{u}_i - \hat{u}_{i-1}}{Q_i - Q_{i-1}}\right) \right) \left(\left(\frac{\hat{u}_{i+1} - \hat{u}_i}{Q_{i+1} - Q_i}\right) + \left(\frac{\hat{u}_i - \hat{u}_{i-1}}{Q_i - Q_{i-1}}\right) \right), (87)$$

$$= -P_i \frac{1}{2} \left(u_x |_{Q_i^-} + u_x |_{Q_i^+} \right), \tag{88}$$

which recovers an equation from the standard (r = 2) Hunter–Saxton case.

Remark 4.9. In the r = 2 case, this formula allows an ad hoc interpretation of mu_x term applied the piecewise linear solutions, when *m* contains Dirac measures centred on points where u_x jumps. This interpretation says that we can just take an average of u_x from each side of the jump point, but this only coincidentally works for the r = 2 case, and is *not true* for r > 2.

Remark 4.10. The same (P, Q) dynamics is obtained from the following variational principle,

$$\delta \hat{S}(\hat{u}, P, Q) = 0, \quad \hat{S} = \int_0^T \hat{l}(\hat{u}, Q) + \sum_{i=1}^n P_i \cdot \left(\dot{Q}_i - \hat{u}_i\right), Q_i(0) = Q_i^A,$$

$$Q_i(T) = Q_i^B, \quad i = 1, \dots, N.$$
(89)

This dynamics has a Q-dependent Lagrangian but trivial Lie algebra action of \hat{u} on Q. The equivalence of these two formulations becomes clear after noticing that they result in the same Hamiltonian after Legendre transformation. However this second variational principle has more directly computable Euler–Lagrange equations, since the vanishing terms above simply do not appear in the first place.

Finally, we examine the $r \to \infty$ limit of equations (76)–(78).

Corollary 4.11. Let $z_i = P^{\frac{1}{r-1}}$. Then, the $r \to \infty$ limit of equations (76)–(78) are given by

$$\dot{Q}_i = \hat{u}_i,\tag{90}$$

$$argmax\left(-\frac{\hat{u}_{i+1}-\hat{u}_i}{Q_{i+1}-Q_i},\frac{\hat{u}_i-\hat{u}_{i-1}}{Q_i-Q_{i-1}}\right) = z_i,\tag{91}$$

$$\dot{z}_i = 0. \tag{92}$$

This limit makes sense, since it keeps \hat{u}_i finite (for distinct Q_i).

Proof. First we examine the difference equation (71). After substituting z_i into P_i and taking the r - 1 root, we obtain

$$z_{i} = \left(-\left| \frac{\hat{u}_{i+1} - \hat{u}_{i}}{Q_{i+1} - Q_{i}} \right|^{r-2} \left(\frac{\hat{u}_{i+1} - \hat{u}_{i}}{Q_{i+1} - Q_{i}} \right) + \left| \frac{\hat{u}_{i} - \hat{u}_{i-1}}{Q_{i} - Q_{i-1}} \right|^{r-2} \left(\frac{\hat{u}_{i} - \hat{u}_{i-1}}{Q_{i} - Q_{i-1}} \right) \right)^{\frac{1}{r-1}}$$
(93)

which recovers (91) in the limit. Substituting z_i into equation (77) gives

$$(r-1)z_i^{r-2}\dot{z}_i = \left(\frac{r-1}{r}\left(\left|\frac{\hat{u}_i - \hat{u}_{i-1}}{Q_i - Q_{i-1}}\right|^r - \left|\frac{\hat{u}_{i+1} - \hat{u}_i}{Q_{i+1} - Q_i}\right|^r\right)\right),\tag{94}$$

and hence

$$\dot{z}_{i} = \frac{1}{(r-1)z_{i}^{r-2}} \left(\frac{r-1}{r} \left(\left| \frac{\hat{u}_{i} - \hat{u}_{i-1}}{Q_{i} - Q_{i-1}} \right|^{r} - \left| \frac{\hat{u}_{i+1} - \hat{u}_{i}}{Q_{i+1} - Q_{i}} \right|^{r} \right) \right),$$
(95)

$$=\frac{1}{(r-1)P_{i}^{\frac{r-2}{r-1}}}\left(\frac{r-1}{r}\left(\left|\frac{\hat{u}_{i}-\hat{u}_{i-1}}{Q_{i}-Q_{i-1}}\right|^{r}-\left|\frac{\hat{u}_{i+1}-\hat{u}_{i}}{Q_{i+1}-Q_{i}}\right|^{r}\right)\right),$$
(96)

$$=\frac{1}{r}\frac{\left(\left|\frac{\hat{u}_{i}-\hat{u}_{i-1}}{Q_{i}-Q_{i-1}}\right|^{r}-\left|\frac{\hat{u}_{i+1}-\hat{u}_{i}}{Q_{i+1}-Q_{i}}\right|^{r}\right)}{\left(-\left|\frac{\hat{u}_{i+1}-\hat{u}_{i}}{Q_{i+1}-Q_{i}}\right|^{r-2}\left(\frac{\hat{u}_{i+1}-\hat{u}_{i}}{Q_{i+1}-Q_{i}}\right)+\left|\frac{\hat{u}_{i}-\hat{u}_{i-1}}{Q_{i}-Q_{i-1}}\right|^{r-2}\left(\frac{\hat{u}_{i}-\hat{u}_{i-1}}{Q_{i}-Q_{i-1}}\right)\right)^{\frac{r-2}{r-1}},$$
(97)

which converges to zero as $r \to \infty$, provided that $(\hat{u}_i - \hat{u}_{i-1})/(Q_i - Q_{i-1})$ is bounded for all *i*.

The well-posedness of equations (76)–(78) depends on the solveability of (78), which is a linear system in the max-algebra (the algebra where the addition operator is replaced by the max operator) that can be formulated as a linear program. The convergence of solutions of equations (76)–(78) to solutions of equations (90) and (91) is an open question, which we investigate through special cases and numerical solutions in section 6.

5. Weak solutions of the *r*-HS equation

As previously discussed, it is natural to hope the results of [CH09, GBR11] allow us to show that the piecewise linear u derived in section 4 satisfies the corresponding Euler–Poincaré equation, i.e., the *r*-Hunter–Saxton equation [14]. However, since the corresponding Lie bracket is not closed in $W^{1,r}$, it is only closed in C^{∞} , the results of those papers do not hold. Indeed, (14) is not well-defined for piecewise linear solutions, not even weakly, since the middle integral in (29),

$$\int_{0}^{1} \left(|u_{x}|^{r-2} u_{x} \right)_{x} w u_{x} \, \mathrm{d}x, \tag{98}$$

is not defined for test functions $w \in H^1([0, 1])$, even after integrating by parts. However, remarkably, the piecewise linear solutions *are* solutions of (32).

Theorem 5.1. *The piecewise linear functions evolving according to the Hamiltonian system above are solutions of equation* (32).

Proof. The proof is via straightforward direct calculation. Let u(x, t) be a time-dependent piecewise linear function over the intervals $[Q_i(t), Q_{i+1}(t)]$, with $u_i = u(Q_i(t))$. Then,

$$\frac{\mathrm{d}}{\mathrm{d}t} \int_{0}^{1} |u_{x}|^{r-2} u_{x} \phi \, \mathrm{d}x = \frac{\mathrm{d}}{\mathrm{d}t} \sum_{i=1}^{n} \int_{Q_{i}}^{Q_{i+1}} \phi \, \mathrm{d}x \left| \frac{u_{i+1} - u_{i}}{Q_{i+1} - Q_{i}} \right|^{r-2} \frac{u_{i+1} - u_{i}}{Q_{i+1} - Q_{i}},$$

$$= \sum_{i=1}^{n} \int_{Q_{i}}^{Q_{i+1}} \phi \, \mathrm{d}x \left| \frac{u_{i+1} - u_{i}}{Q_{i+1} - Q_{i}} \right|^{r-2} \frac{u_{i+1} - u_{i}}{Q_{i+1} - Q_{i}}$$

$$+ \sum_{i=1}^{n} (\dot{Q}_{i+1}\phi(Q_{i+1}) - \dot{Q}_{i}\phi(Q_{i})) \left| \frac{u_{i+1} - u_{i}}{Q_{i+1} - Q_{i}} \right|^{r-2} \frac{u_{i+1} - u_{i}}{Q_{i+1} - Q_{i}}.$$
(99)

Now, since *u* is piecewise linear, u_x is constant in every interval $[Q_i, Q_{i+1}]$, hence

$$\left(u|u_x|^{r-2}u_x\right)_x = |u_x|^r,$$
(100)

and, upon splitting the integral into intervals and integrating by parts,

$$-\int_{0}^{1} u |u_{x}|^{r-2} u_{x} \phi_{x} \, \mathrm{d}x = -\sum_{i=0}^{n} \int_{Q_{i}}^{Q_{i+1}} u |u_{x}|^{r-2} u_{x} \phi_{x} \, \mathrm{d}x,$$

$$= -\sum_{i=0}^{n} \left(u_{i+1} \phi(Q_{i+1}) - u_{i} \phi(Q_{i}) \right) \left| \frac{u_{i+1} - u_{i}}{Q_{i+1} - Q_{i}} \right|^{r-2} \frac{u_{i+1} - u_{i}}{Q_{i+1} - Q_{i}}$$

$$+ \sum_{i=1}^{n} \int_{Q_{i}}^{Q_{i+1}} |u_{x}|^{r} \phi \, \mathrm{d}x,$$

$$= -\sum_{i=0}^{n} \left(u_{i+1} \phi(Q_{i+1}) - u_{i} \phi(Q_{i}) \right) \left| \frac{u_{i+1} - u_{i}}{Q_{i+1} - Q_{i}} \right|^{r-2} \frac{u_{i+1} - u_{i}}{Q_{i+1} - Q_{i}}$$

$$+ \sum_{i=1}^{n} \int_{Q_{i}}^{Q_{i+1}} \phi \, \mathrm{d}x \left| \frac{u_{i+1} - u_{i}}{Q_{i+1} - Q_{i}} \right|^{r}.$$
(101)

Combining these together, we have

$$c(t)\sum_{i=1}^{n} \Phi_{i} = \sum_{i=1}^{n} \left((\dot{Q}_{i+1} - u_{i+1})\phi(Q_{i+1}) - (\dot{Q}_{i} - u_{i})\phi(Q_{i}) \right) \\ \times \left| \frac{u_{i+1} - u_{i}}{Q_{i+1} - Q_{i}} \right|^{r-2} \frac{u_{i+1} - u_{i}}{Q_{i+1} - Q_{i}} + \sum_{i=1}^{n} \Phi_{i} \frac{d}{dt} \left| \frac{u_{i+1} - u_{i}}{Q_{i+1} - Q_{i}} \right|^{r-2} \frac{u_{i+1} - u_{i}}{Q_{i+1} - Q_{i}} \\ + \sum_{i=1}^{n} \Phi_{i} \frac{r-1}{r} \left| \frac{u_{i+1} - u_{i}}{Q_{i+1} - Q_{i}} \right|^{r},$$
(102)

where

$$\Phi_i = \int_{Q_i}^{Q_{i+1}} \phi \, \mathrm{d}x, \quad i = 1, \dots, n.$$
(103)

Taking $\dot{Q}_i = u_i$, we are left with

$$\frac{\mathrm{d}}{\mathrm{d}t} \left| \frac{u_{i+1} - u_i}{Q_{i+1} - Q_i} \right|^{r-2} \frac{u_{i+1} - u_i}{Q_{i+1} - Q_i} + \frac{r-1}{r} \left| \frac{u_{i+1} - u_i}{Q_{i+1} - Q_i} \right|^r = c(t),$$
(104)

since Φ_i are arbitrary. Now we show that our singular solutions also satisfy this equation. First, define c(t) according to

$$c(t) = \frac{\mathrm{d}}{\mathrm{d}t} \left| \frac{u_1 - u_0}{Q_1 - Q_0} \right|^{r-2} \frac{u_1 - u_0}{Q_1 - Q_0} + \frac{r - 1}{r} \left| \frac{u_1 - u_0}{Q_1 - Q_0} \right|^r.$$
(105)

Now combine equations (71) and (77), we have

$$0 = \frac{d}{dt} \left(-\left| \frac{u_{i+1} - u_i}{Q_{i+1} - Q_i} \right|^{r-2} \frac{u_{i+1} - u_i}{Q_{i+1} - Q_i} + \left| \frac{u_i - u_{i-1}}{Q_i - Q_{i-1}} \right|^{r-2} \frac{u_i - u_{i-1}}{Q_i - Q_{i-1}} \right) + \frac{r - 1}{r} \left(-\left| \frac{\hat{u}_i - \hat{u}_{i-1}}{Q_i - Q_{i-1}} \right|^r + \left| \frac{\hat{u}_{i+1} - \hat{u}_i}{Q_{i+1} - Q_i} \right|^r \right),$$
(106)

which means that

$$\frac{\mathrm{d}}{\mathrm{d}t} \left(\left| \frac{u_{i+1} - u_i}{Q_{i+1} - Q_i} \right|^{r-2} \frac{u_{i+1} - u_i}{Q_{i+1} - Q_i} \right) + \frac{r-1}{r} \left| \frac{u_{i+1} - u_i}{Q_{i+1} - Q_i} \right|^r$$
$$= \frac{\mathrm{d}}{\mathrm{d}t} \left(\left| \frac{u_i - u_{i-1}}{Q_i - Q_{i-1}} \right|^{r-2} \frac{u_i - u_{i-1}}{Q_i - Q_{i-1}} \right) + \frac{r-1}{r} \left| \frac{u_i - u_{i-1}}{Q_i - Q_{i-1}} \right|^r, \tag{107}$$

and hence

$$\frac{\mathrm{d}}{\mathrm{d}t} \left(\left| \frac{u_{i+1} - u_i}{Q_{i+1} - Q_i} \right|^{r-2} \frac{u_{i+1} - u_i}{Q_{i+1} - Q_i} \right) + \frac{r-1}{r} \left| \frac{u_{i+1} - u_i}{Q_{i+1} - Q_i} \right|^r = c(t),$$
(108)

by induction, and we have our required equation.

6. Numerical examples

In this section we compute a number of examples of piecewise linear solutions, obtained by integrating equations (76)-(78).

6.1. One point solutions

In the case of a single point Q_1 , we may use energy conservation to find an algebraic dependence between \hat{u}_1 and Q_1 . Specialising to [a, b] = [0, 1] to simplify, the conserved energy is

$$E = \frac{1}{r} \left(Q_1 \left(\frac{\hat{u}_1}{Q_1} \right)^r + (1 - Q_1) \left(\frac{\hat{u}_1}{1 - Q_1} \right)^r \right) = \frac{\hat{u}_1^r}{r} \left(Q_1^{1-r} + (1 - Q_1)^{1-r} \right).$$
(109)

Since E is time-independent, we may compute it from the initial condition, and then

$$\hat{u}_1 = \left(\frac{rE}{Q_1^{1-r} + (1-Q_1)^{1-r}}\right)^{\frac{1}{r}},\tag{110}$$

assuming a positive root (it is a consequence of the one point *P* equation that *u* will stay positive if it is positive initially). We may then integrate the first-order equation $\dot{Q}_1 = \hat{u}_1(Q_1)$ (this must be done numerically in general). Since *u* is always positive, Q_1 is monotonically increasing,

$$\hat{u}_1 \sim E^{1/r} r^{1/r} \epsilon^{(r-1)/r}, \quad \text{as } \epsilon \to 0.$$
 (111)

The equation

$$\frac{\mathrm{d}\epsilon}{\mathrm{d}t} = -(Er)^{1/r} \epsilon^{(r-1)/r} \tag{112}$$

has solution

$$\epsilon = \left(\epsilon_0^{1/r} - \frac{(Er)^{1/r}t}{r}\right)^r,\tag{113}$$

which converges to zero in finite time. At the same time, the derivative of the piecewise linear function u in the region $[Q_1, 1]$ is given by

$$u_x = \frac{-\hat{u}_1}{1 - Q_1} \sim (Er)^{1/r} \epsilon^{-1/r} \to \infty \quad \text{as } \epsilon \to 0,$$
(114)

indicating that the $W^{1,\infty}$ -norm blows up in finite time.

Snapshots of a typical solution are shown in figure 1. A comparison of the behaviour of single point solutions is made in figure 2. We observe that the $\hat{u}_1 - Q_1$ relation approaches $\hat{u}_1 = \min(Q_1, 1 - Q_1)$ as $r \to \infty$. A plot of the derivative u_x in $[Q_1, 1]$ is given in figure 3. We observe that whilst the solution always blows up in finite time, the blow up time is later for higher r, and is roughly proportional to r, as predicted by the asymptotic approximation. This is also compatible with the intuition that the $r \to \infty$ limit should preserve the $W^{1,\infty}$ norm.

For the single point case, the $r \to \infty$ limit suggested in corollary 4.11 results in the equations

$$z_1 = \max\left(\frac{\hat{u}_1}{Q_1}, \frac{\hat{u}_1}{1 - Q_1}\right), \quad \dot{z}_1 = 0.$$
 (115)

We also deduce that $\dot{z}_1 = 0$ from taking the limit in the P_1 -equation. For initial conditions $Q_1(0) > 0.5$, we obtain

$$Q_1(t) = 1 - (1 - Q_1(0))e^{-z_1 t},$$
(116)

for z_1 constant and positive. This means that z_1 does not reach 1 in finite time, and hence the solution does not blow up in finite time. The infinite limit solutions are compared with the finite *r* solutions in figure 4.

6.2. Two point solutions

For two or more points, unless a symmetric solution is sought, there is no closed form for the $\hat{u}_1 - Q_1$ relation and equation (78) must be solved iteratively using Newton's method. The result can then be used in an explicit time-integrator such as the standard 4th-order Runge–Kutta scheme that we used for these examples. In this section, we consider three cases, as follows.

- (a) Symmetric collision: $(Q_1, \hat{u}_1) = (0.1, 0.1), (Q_2, \hat{u}_2) = (0.9, -0.1),$
- (b) Chasing collision: $(Q_1, \hat{u}_1) = (0.1, 0.2), (Q_2, \hat{u}_2) = (0.2, 0.1)$, and
- (c) Asymmetric collision: $(Q_1, \hat{u}_1) = (0.1, 0.2), (Q_2, \hat{u}_2) = (0.2, -0.125).$



Figure 1. Plots of u(x) for a one point solution with r = 2, at times t = 0., 1.83, 3.67, 5.5.



Figure 2. Comparison of one point trajectories for different values of r, 2, 4, 6, 8, 10, 12, 14, 16, 18, 20, with initial condition $(Q_1, \hat{u}_1) = (0.1, 0.1)$. Left: the point locations $q = Q_1(t)$ are plotted against time. Right: the peak velocity \hat{u}_1 is plotted against $q = Q_1(t)$.

The numerical solutions are shown in figure 5. We note that finite time singularities occur in all three cases, even the chasing collision (in contrast to the same case for the Camassa–Holm equation, where the solitons transfer momentum at a distance and the solution does not blow up). Finally, we investigate the approximation of smooth solutions by piecewise-linear solutions and their subsequent evolution in time. We take 101 equispaced points on the interval



Figure 3. Plot of $|u_x|$ in the region $[Q_1, 1]$, where it is independent of *x* and denoted $|u_x^+|$, against time for various *r*, with the same initial condition as figure 2, $(Q_1, \hat{u}) = (0.1, 0.1)$. The time to singularity increases with *r*.



Figure 4. Comparison of $r \to \infty$ limiting solutions with finite *r* solutions, as per figure 2.

[0, 1] and interpolate the function $u(x) = \sin(2\pi x)$ to produce an initial condition for $\{Q_i\}_{i=0}^{100}$, $\{P_i\}_{i=0}^{100}$. We observe a strong jump in the derivative emerging for large *r*. The results are shown in figure 6.

7. Summary and outlook

In this paper we introduced an extension of the Hunter–Saxton equation, which we call the r-Hunter–Saxton equation. These equations are the Euler–Poincaré equations with reduced Lagrangian given by the $W^{1,r}$ norm, and are associated with the evolution of geodesics in the diffeomorphism group with metric defined by the $W^{1,r}$ norm. This replaces the linear Laplacian relating momentum and velocity by a nonlinear r-Laplacian. We introduced an optimal control variational principle for parameterising finite dimensional subspaces of the solutions of



Figure 5. Two point solutions. For each r value, the trajectories for both points $(u, q) = (\hat{u}_i, Q_i)$ are shown in the same colour. Top: symmetric collision. The points have equal velocities with opposite sign, and are on course to a head-on collision. On the left are (q, t) plots, with $q = Q_i$ for i = 1, 2, showing collisions at progressively later times for larger r. On the right are phase plane plots with $u = \hat{u}_i(t)$ plotted versus $q = Q_i(t)$ We see that the velocity starts decaying earlier for larger r, but not decaying to zero so a collision always occurs in finite time. Middle: chasing collision. Both points are moving to the right, but the one behind is moving faster, and eventually catches up in a collision. Bottom: asymmetric collision. The points have velocities of opposite sign, with the point on the left moving faster, so the solution is not symmetric. In this solution, a new feature emerges for r > 2, with the initially left-moving slower point on the right eventually moving right before eventual collision. This contrasts with the r = 2 case, where the left-moving point does not undergo a reversal.



Figure 6. Evolution of piecewise linear solutions u(x) initialised from interpolation of smooth functions, shown at times t = 0, 1.33, 2, 67, 4. A strong jump in the derivative is emerging for large *r*.

PDEs, and derived a Hamiltonian system for a finite dimensional set of points Q_1, \ldots, Q_n plus their conjugate momenta. This corresponds to piecewise linear functions u with jumps in the derivative at each of these points. Although these piecewise functions are not smooth enough to satisfy the *r*-Hunter–Saxton equation, we showed that remarkably they are still weak solutions of an integrated form of the *r*-Hunter–Saxton equation.

There are plenty of interesting open questions about the *r*-Hunter–Saxton equations and the piecewise linear solutions in the limit as $r \to \infty$. We have presented some evidence that solutions of equations (76)–(78) converge to solutions of equations (90) and (91) (before blow-up). We have also presented evidence that solutions of (90) and (91) exist for all times, whilst solutions of (76)–(78) blow up in finite time.

It is very interesting to link the solutions to (90) and (91) back to the $r \to \infty$ limit of the optimal control problem in definition 4.1. This is dangerous, since it involves exchanging the limits $r \to \infty$ and the limits defining variational derivatives. However, it is tantalising that the solutions to (90) and (91) preserve their $W^{1,\infty}$ norm.

Another interesting question is whether solutions of (76)–(78) can be used to approximate smooth solutions of the *r*-Hunter–Saxton equation. The control provided over the $W^{1,p}$ norm of the numerical solution should provide a useful tool for this.

A final direction of enquiry is, what is the correct $r \to \infty$ limit of the *r*-Hunter–Saxton equation? Can the solutions of the limiting equation be approximated by (90) and (91)? Does the limiting solution preserve the $W^{1,\infty}$ norm? If so this would provide a way to establish existence for arbitrary time intervals. Do the limiting solutions generate geodesics for diffeomorphisms in 1D under the $W^{1,\infty}$ -norm? A useful starting point in this direction might be the

recent work of [BM19], who extended the work of [Len07] to $W^{1,r}$ metrics through mappings to the L^r -sphere, where the $r \to \infty$ limit may be treated more easily.

All of these intriguing questions will be the subject of future work.

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